



Screening binary redox pairs for solar thermochemical ammonia synthesis using machine learned predictions of the Gibbs formation energies at finite temperatures

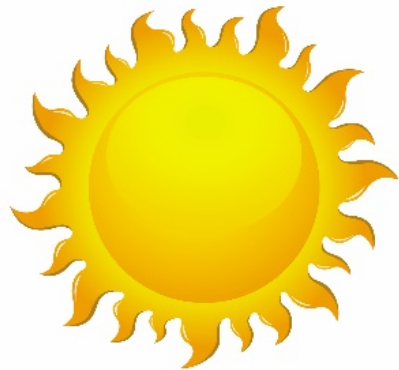
Christopher J. Bartel

John R. Rumptz, Aaron M. Holder, Alan W. Weimer, Charles B. Musgrave

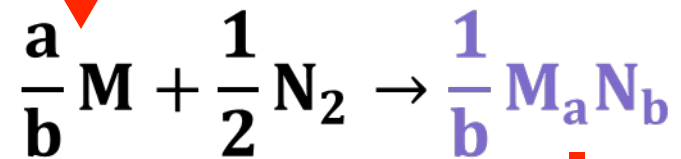
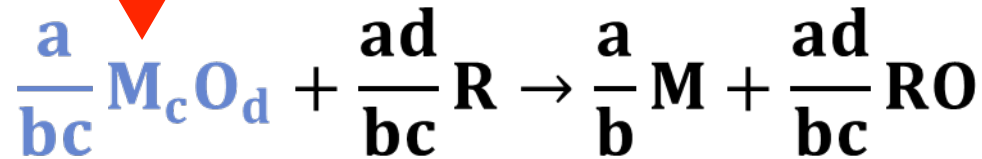
AIChE Annual Meeting

Minneapolis, MN Nov 2, 2017

Solar thermochemical NH₃ synthesis (STAS)



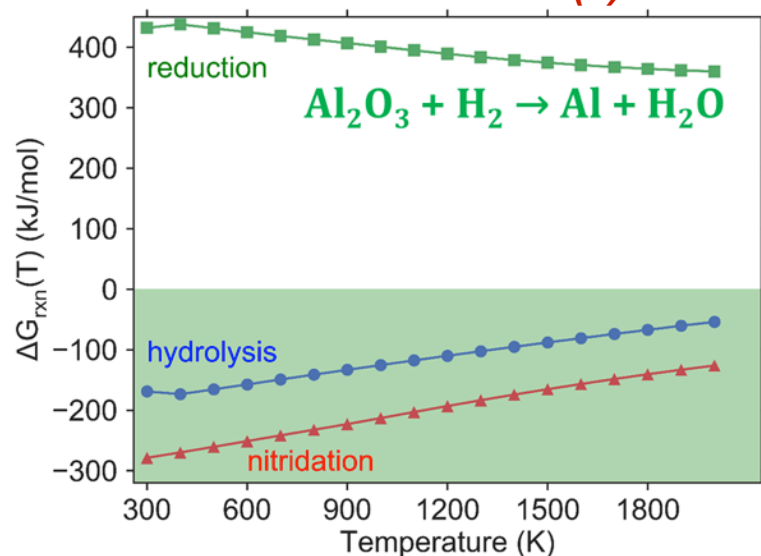
Reduction/oxidation of
metal oxide/metal nitride
pair driven by concentrated
solar radiation



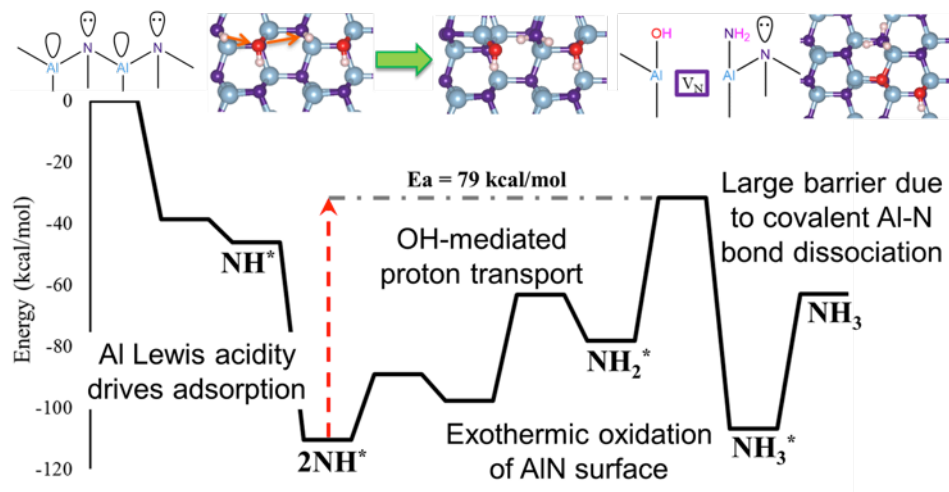
**AlN/Al₂O₃ the first
proposed redox pair**

AlN/Al₂O₃ not the ideal redox pair

Reduction of Al₂O₃ requires C_(s)



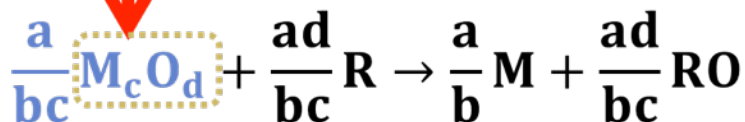
Kinetics of hydrolysis are challenging



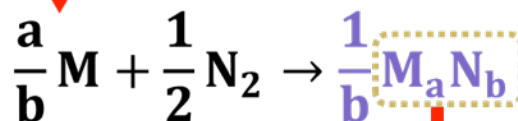
Identifying viable pairs (MN/MO)



Reduction/oxidation of
metal oxide/metal nitride
pair driven by concentrated
solar radiation



*What about other
materials?!*



What makes a good pair (based on thermo)?

- Rxns are spontaneous and high yield
- MN and MO are thermodynamically stable

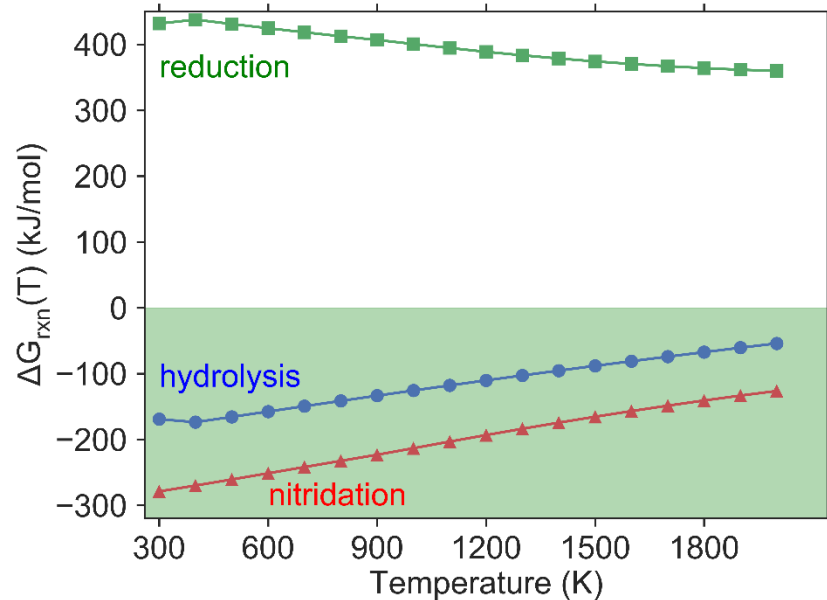
Identifying viable pairs (MN/MO)

What makes a good pair?

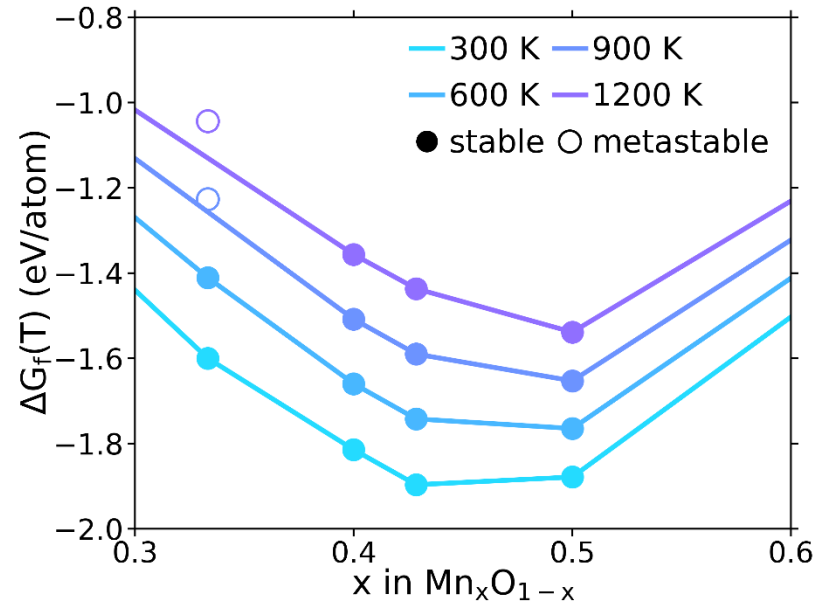
- Rxns are spontaneous and high yield
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$$\Delta G_f(T)$$

$$\Delta G_{\text{rxn}}(T) < 0$$



$$\Delta G_d(T) < 0$$



Acquiring Gibbs formation energies

Experimental data is
incomplete (and difficult)



~100s of experimental $\Delta G_f(T)$
(~20 MN)

Aluminum Oxide, Alpha (Al_2O_3) $\text{Al}_2\text{O}_3(\text{cr})$

Enthalpy Reference Temperature = $T_r = 298.15 \text{ K}$

Standard State Pressure = $p^\circ = 0.1 \text{ MPa}$

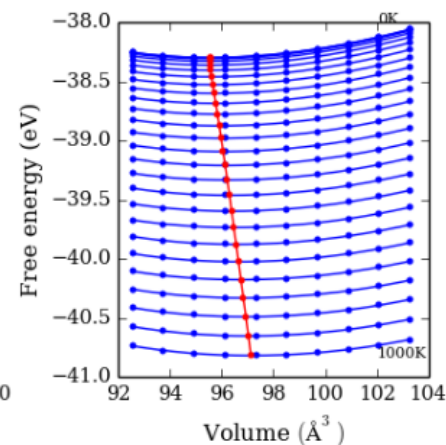
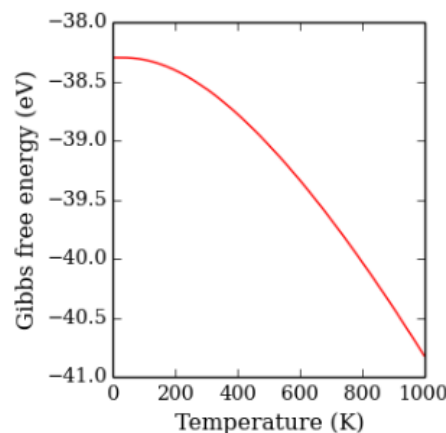
T/K	$\text{J}\cdot\text{K}^{-1}\cdot\text{mol}^{-1}$			$\text{kJ}\cdot\text{mol}^{-1}$			
	C_p°	S°	$-\left[G^\circ - H^\circ(T_r)\right]/T$	$H - H^\circ(T_r)$	$\Delta_f H^\circ$	$\Delta_f G^\circ$	$\log K_f$
0	0.	0.	INFINITE	-10.020	-1663.608	-1663.608	INFINITE
100	12.855	4.295	101.230	-9.693	-1668.606	-1641.642	857.506
200	51.120	24.880	57.381	-6.500	-1673.383	-1612.656	421.183
298.15	79.015	50.950	50.950	0.	-1675.692	-1582.275	277.208
300	79.416	51.440	50.951	0.147	-1675.717	-1581.696	275.398
400	96.086	76.779	54.293	8.995	-1676.342	-1550.226	202.439
500	106.131	99.388	61.098	19.145	-1676.045	-1518.718	158.659

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phonondb

~100s of calculated $G(T)$



Acquiring Gibbs formation energies

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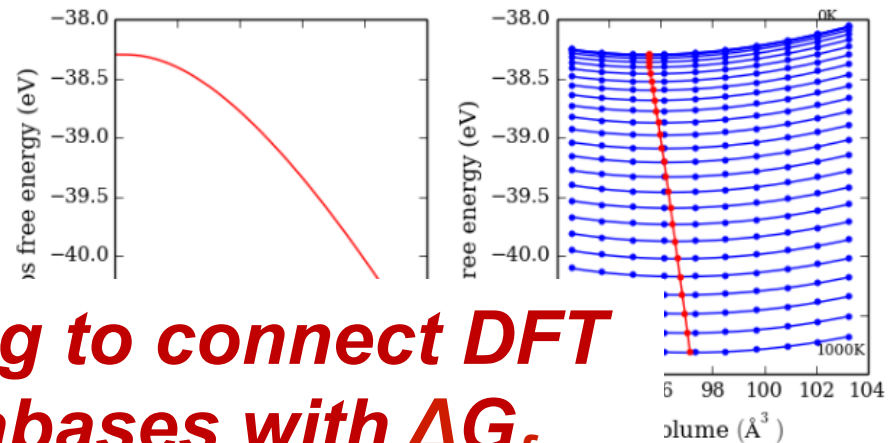
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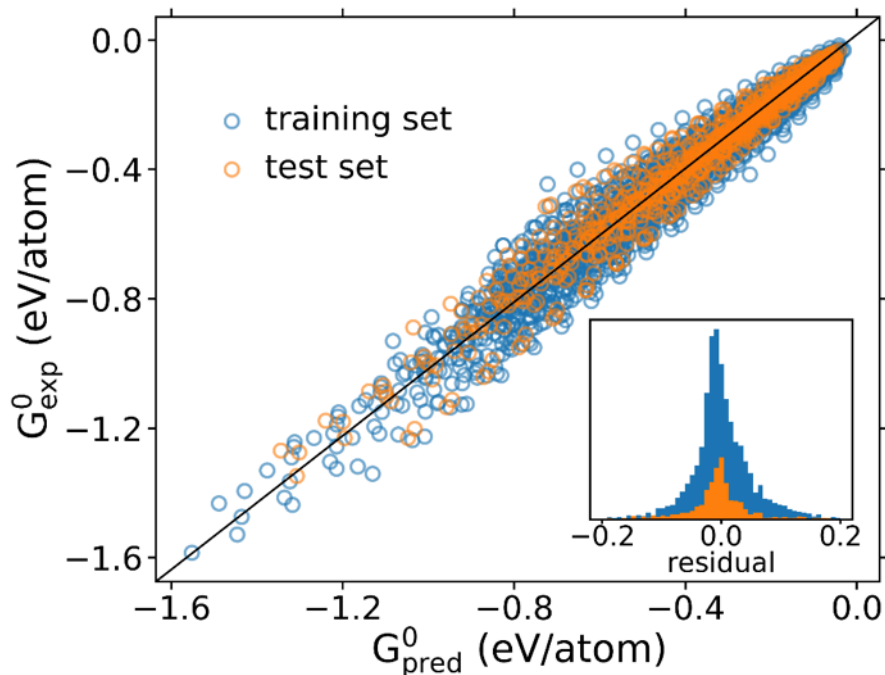
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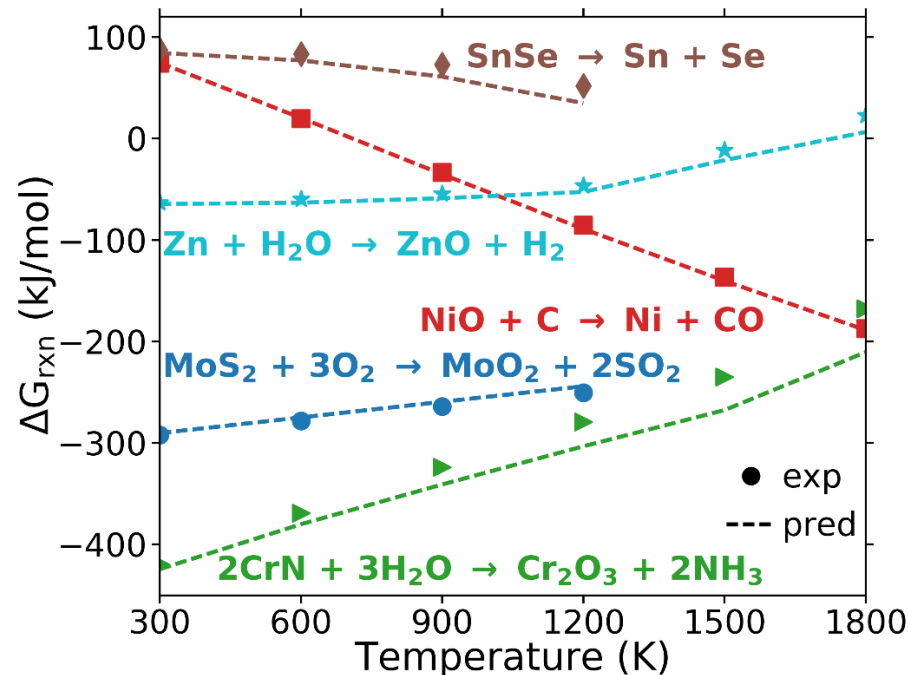
**Machine learning to connect DFT
materials databases with ΔG_f**

Some evidence that ML model works

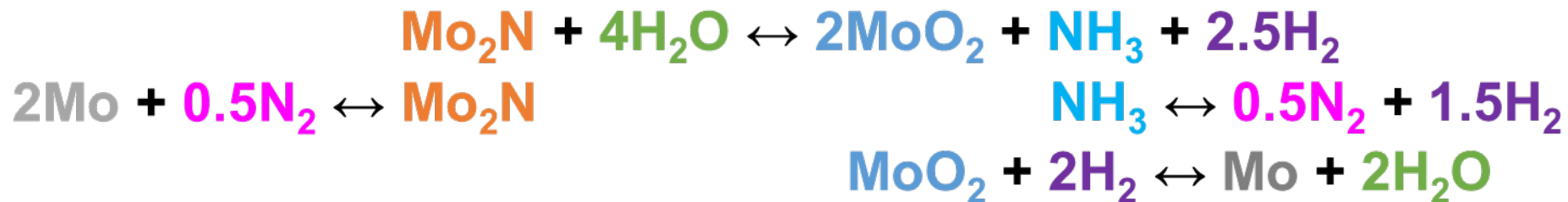
ML model performs similarly on training and test sets



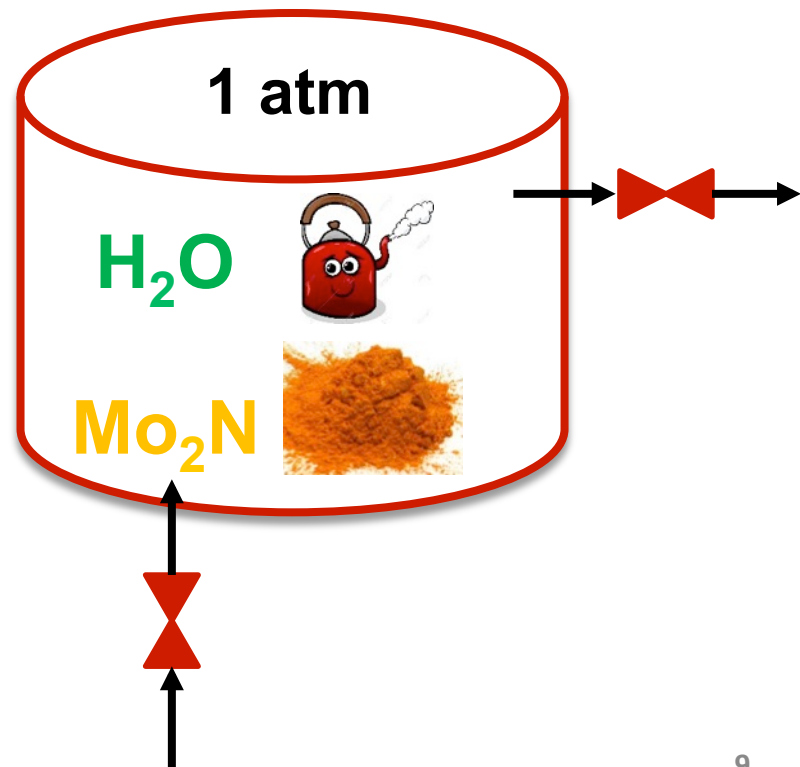
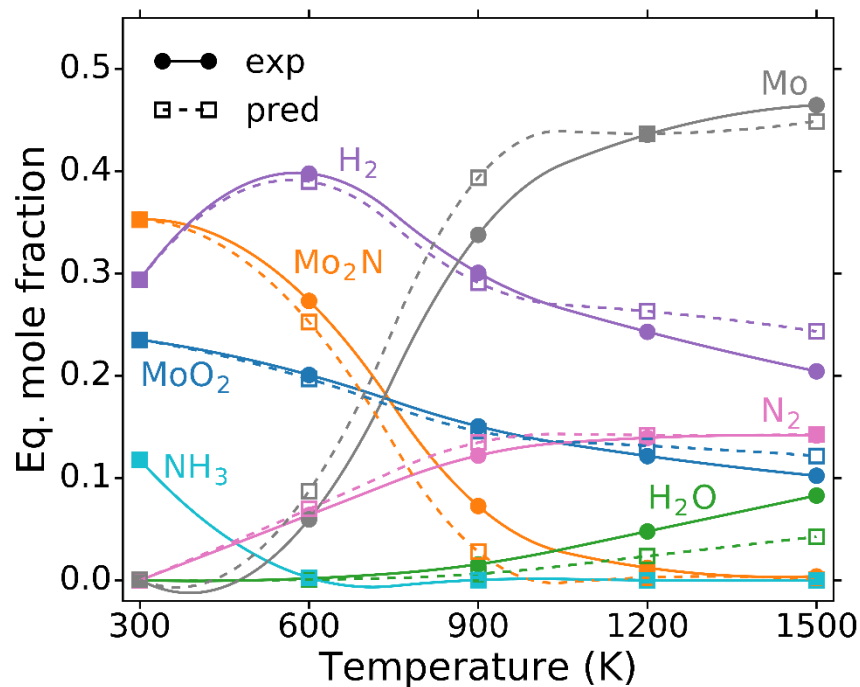
ML model captures reaction energetics well



Some evidence that ML model works



Complex thermodynamic equilibrium well predicted



Identifying viable pairs (MN/MO)

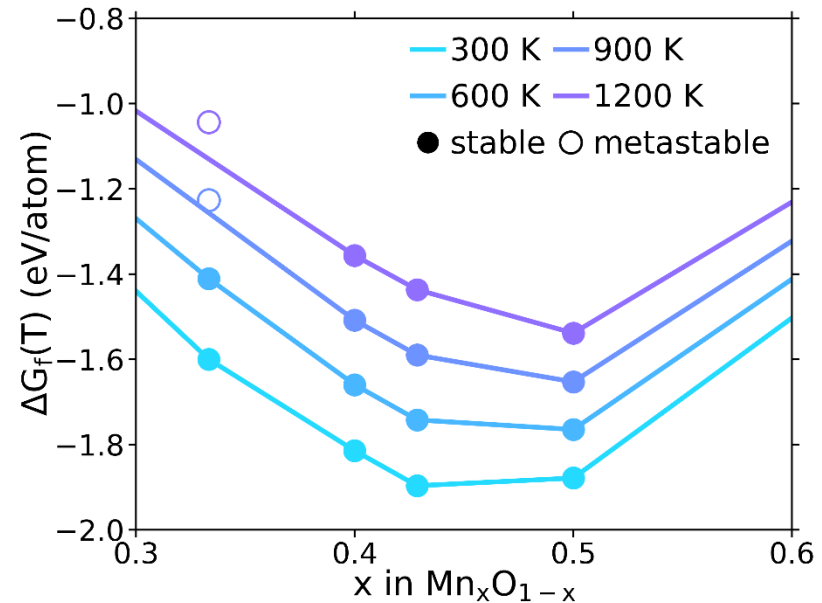
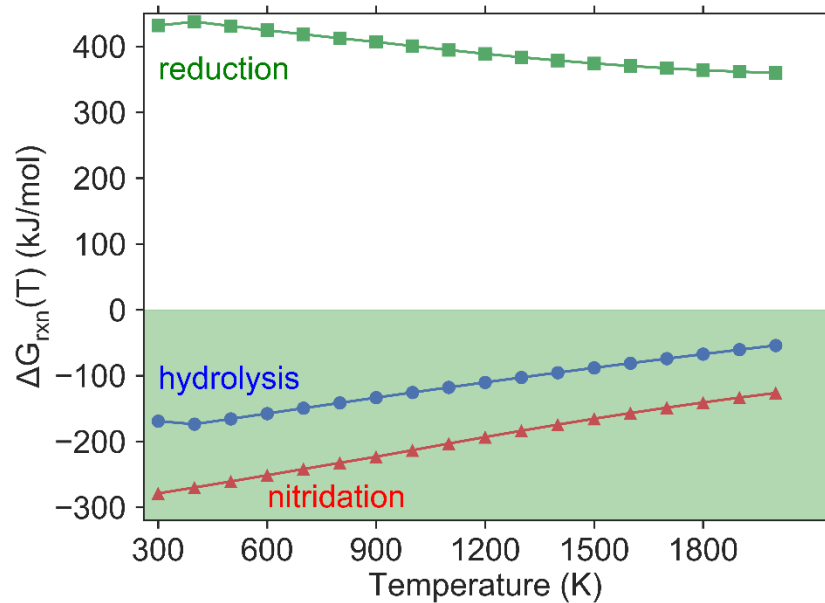
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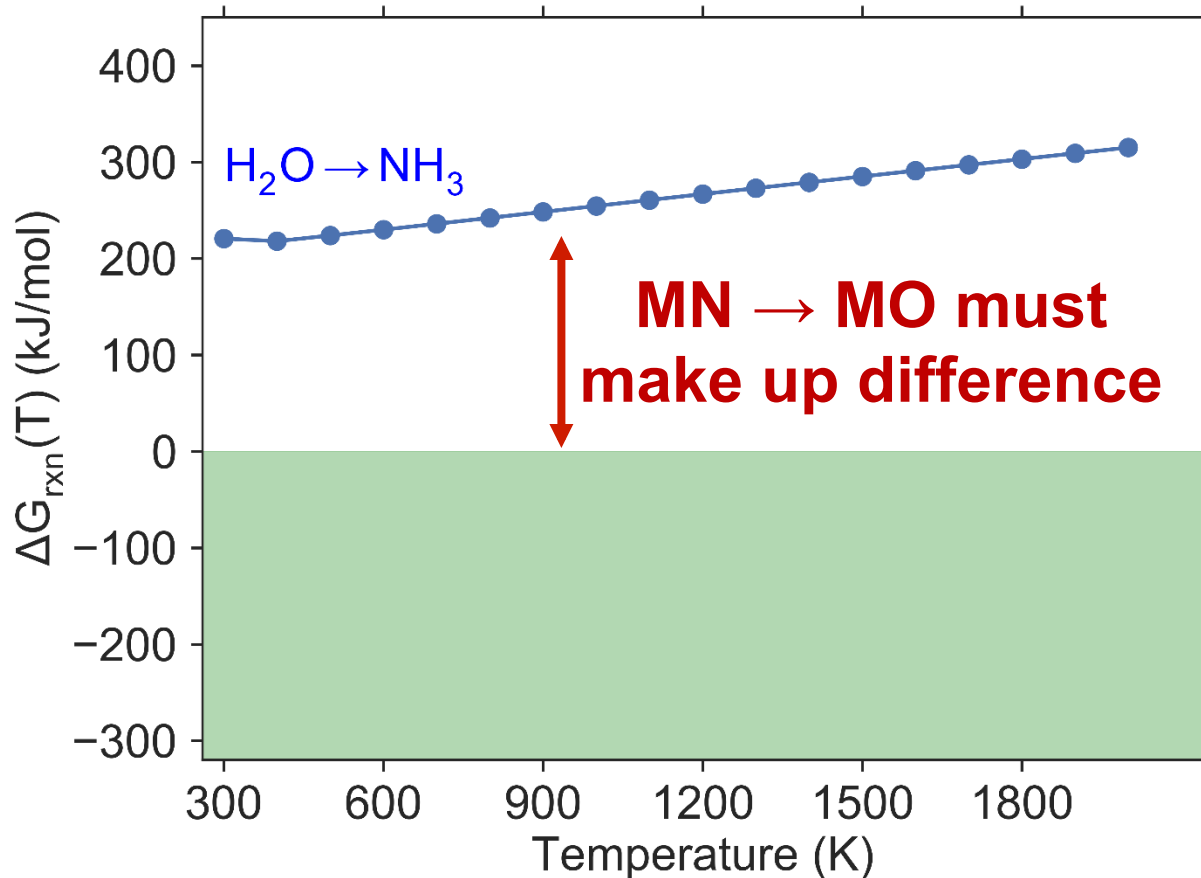
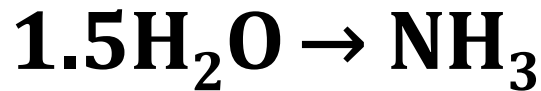
$$\Delta G_d(T) < 0$$



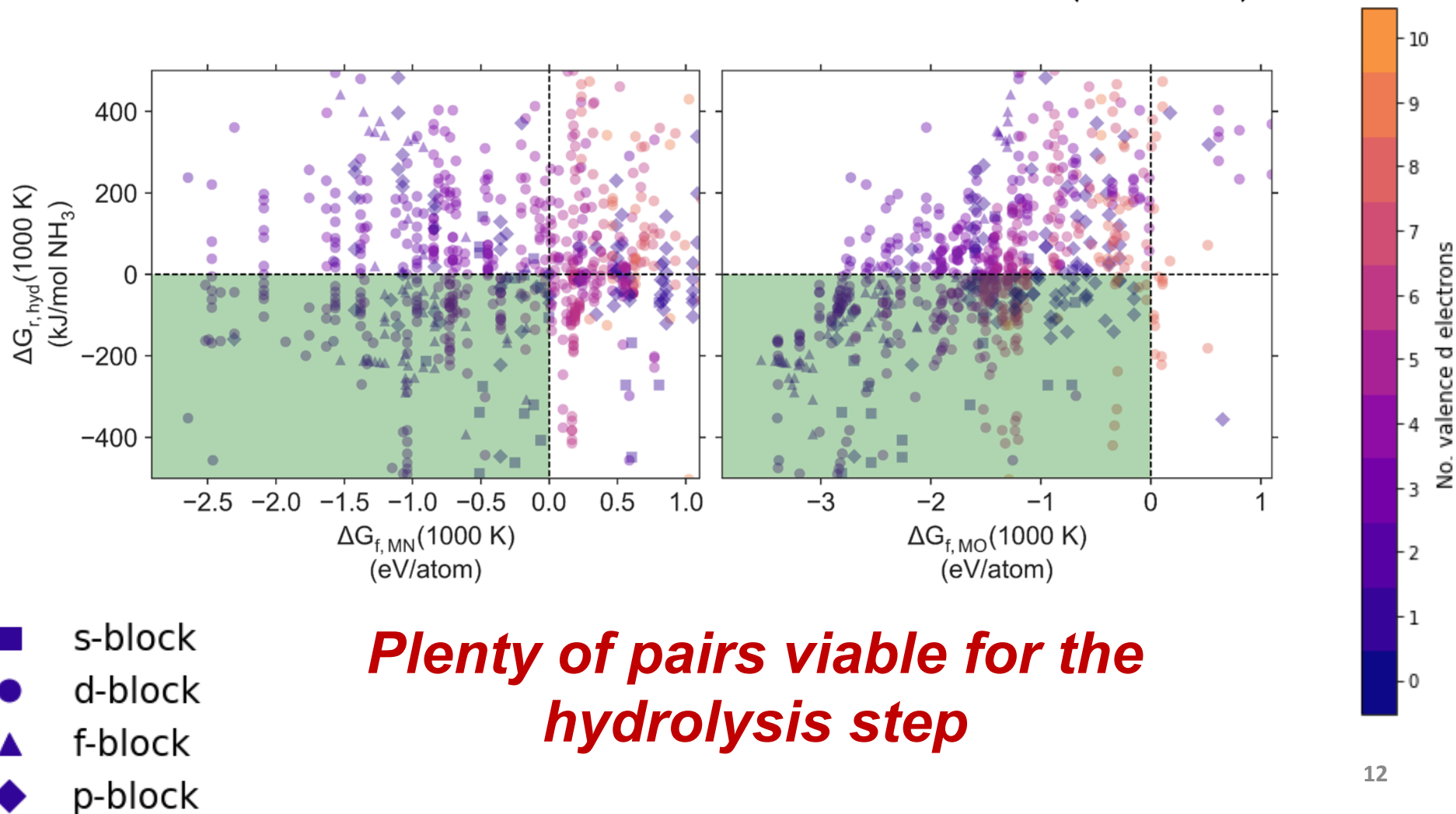
**MATERIALS
PROJECT**



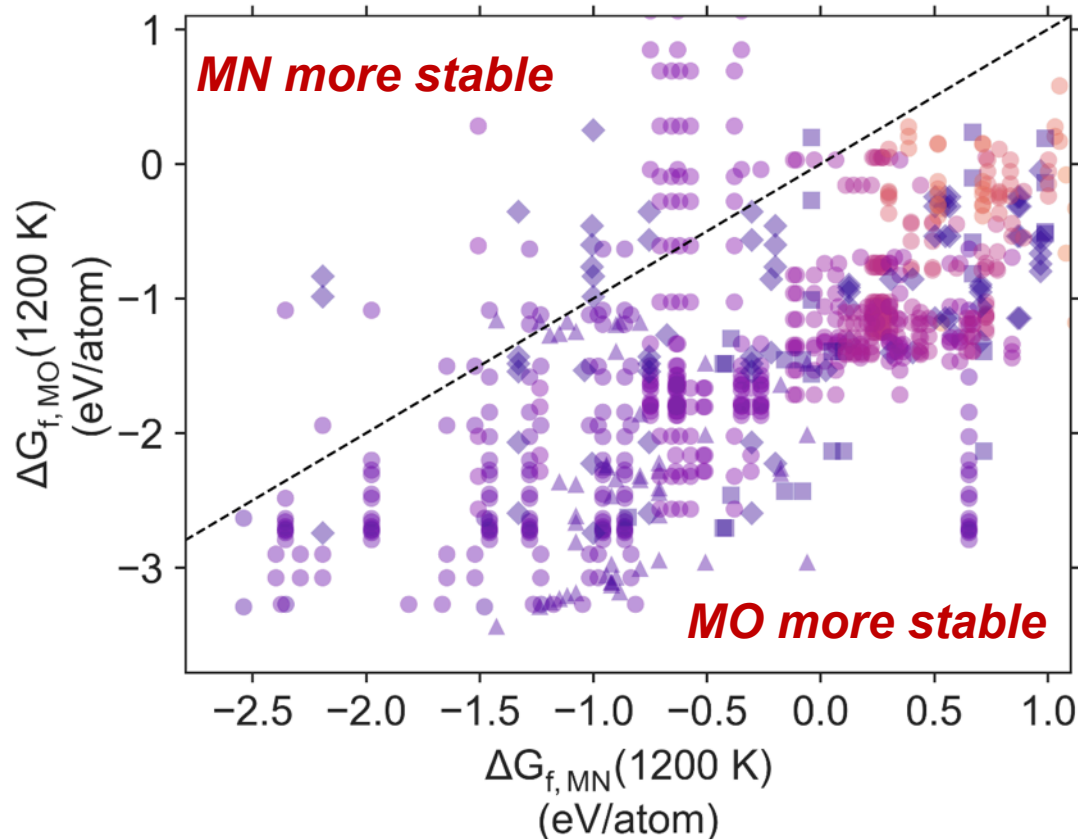
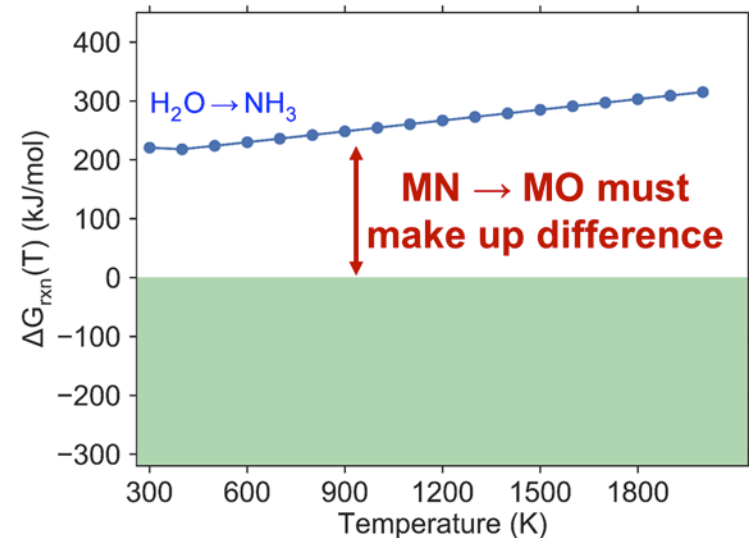
Viability of pairs for NH_3 generation



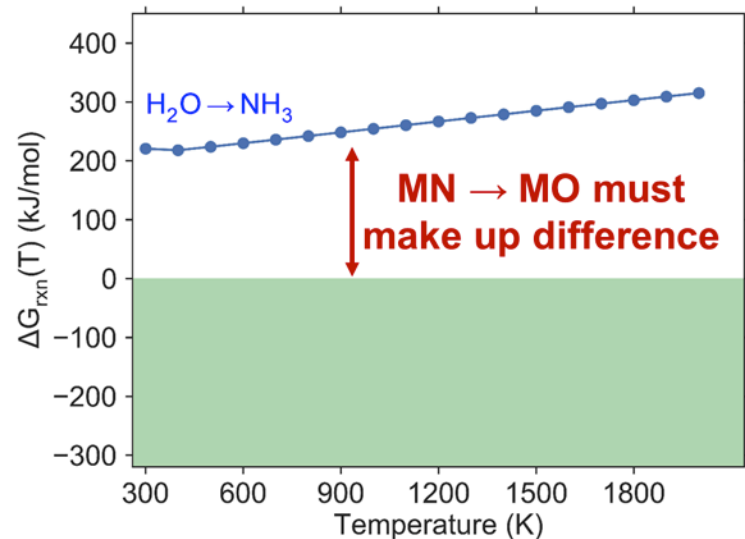
Viability of pairs for NH₃ generation



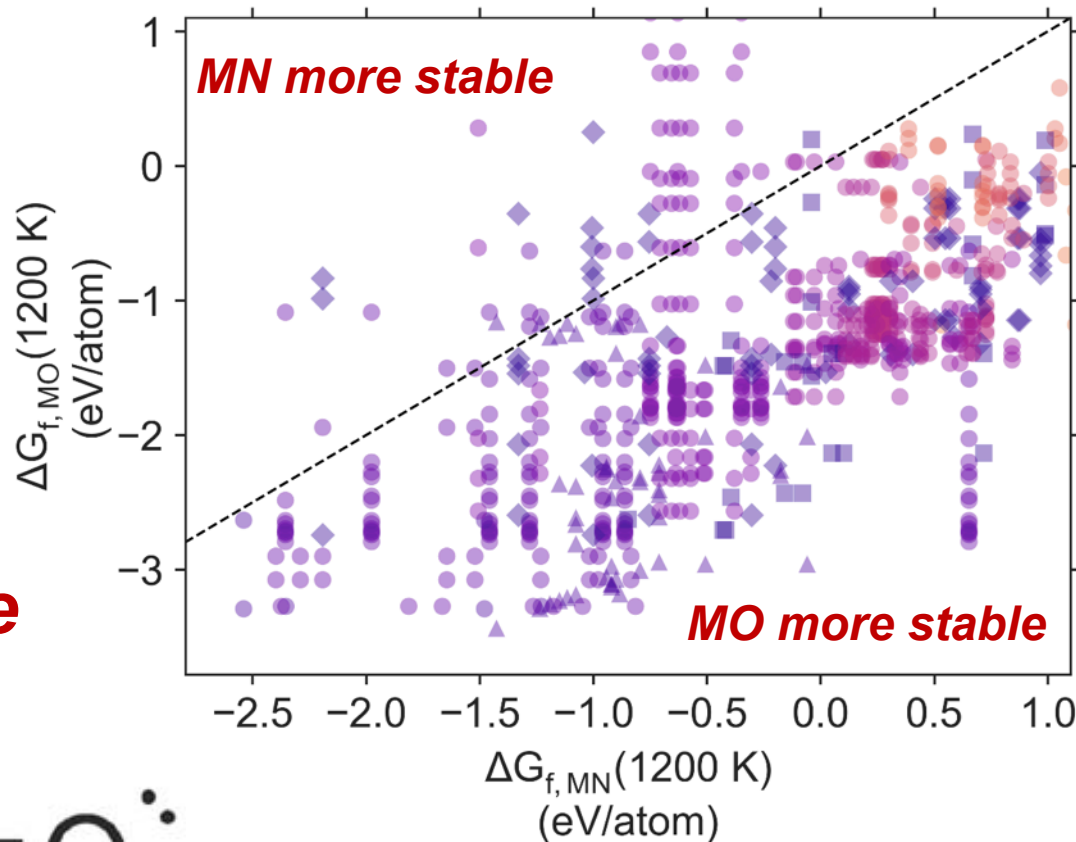
Viability of pairs for NH_3 generation



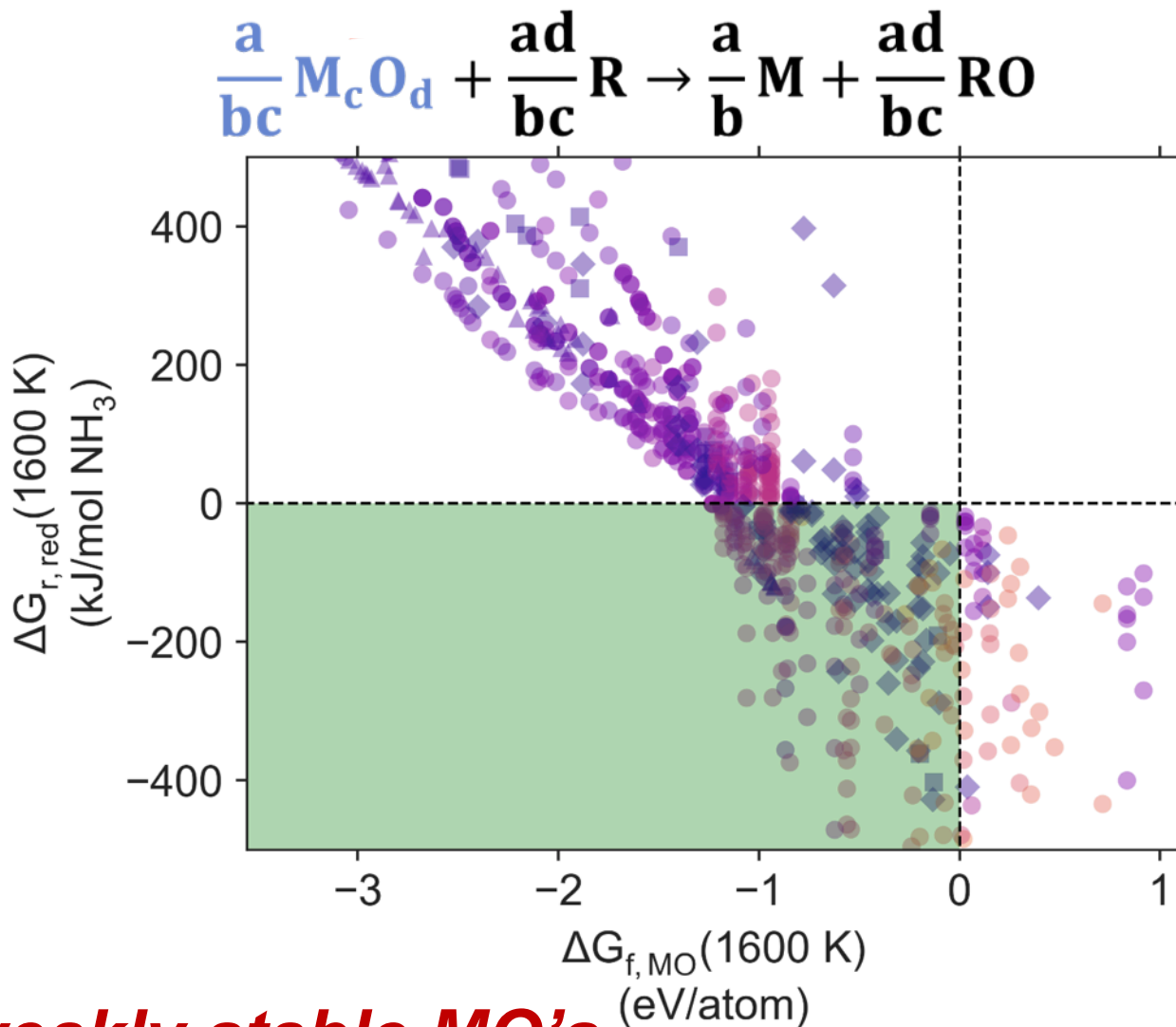
Viability of pairs for NH_3 generation



In general, MO more stable than MN

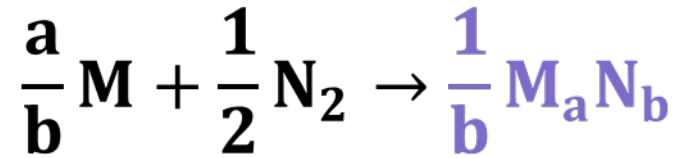


Viability of pairs for MO reduction by H_2

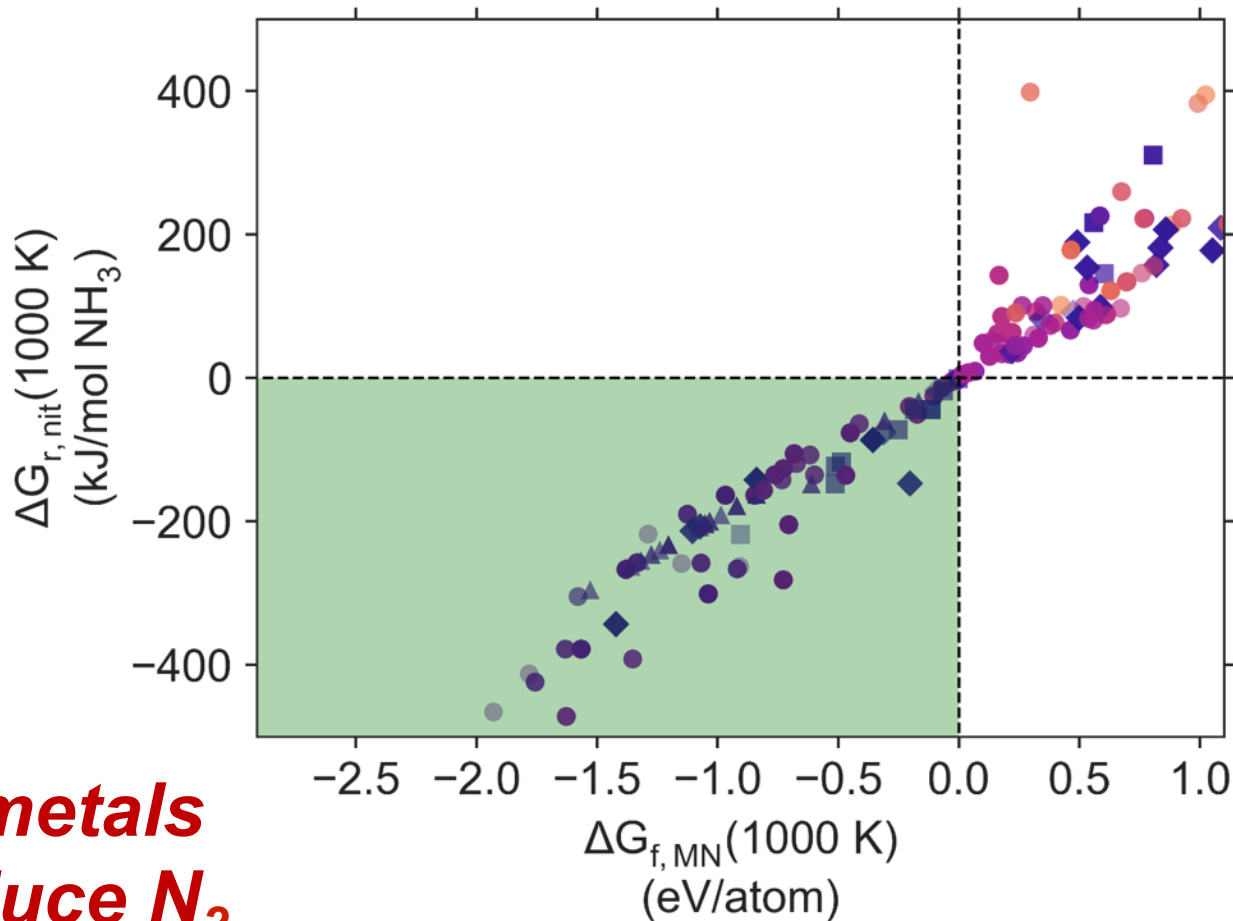


**Many weakly stable MO's
are reducible at 1600 K**

Viability of pairs for N fixation

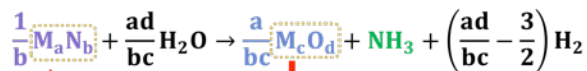


$$\Delta G_{f,MN} < 0$$

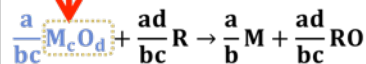


*Many metals
can reduce N_2*

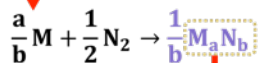
But what about the full cycle?



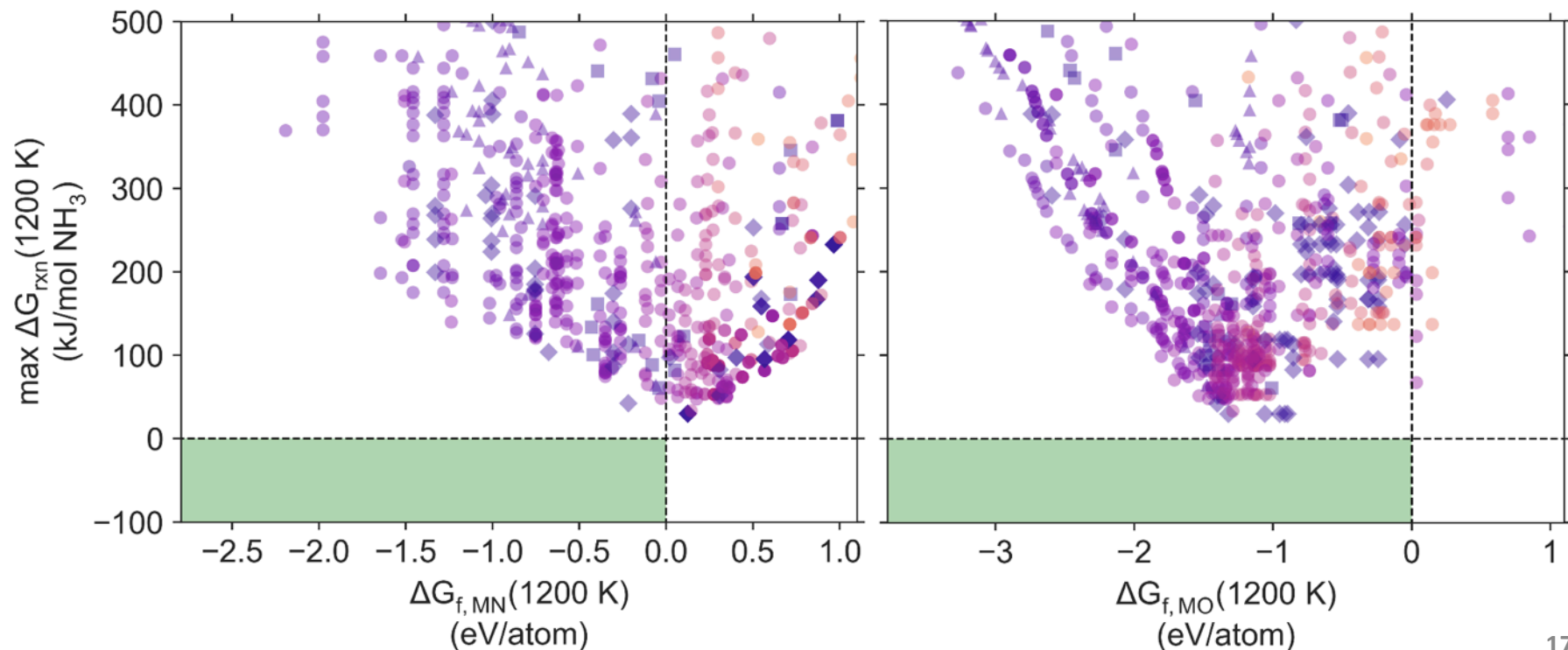
Reduction/oxidation of metal oxide/metal nitride pair driven by concentrated solar radiation



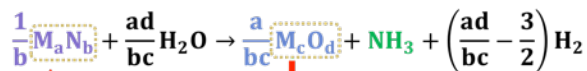
What about other materials?!



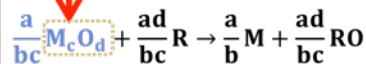
**Correlated stability of MN/
MO limits viability**



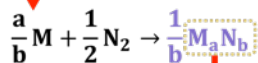
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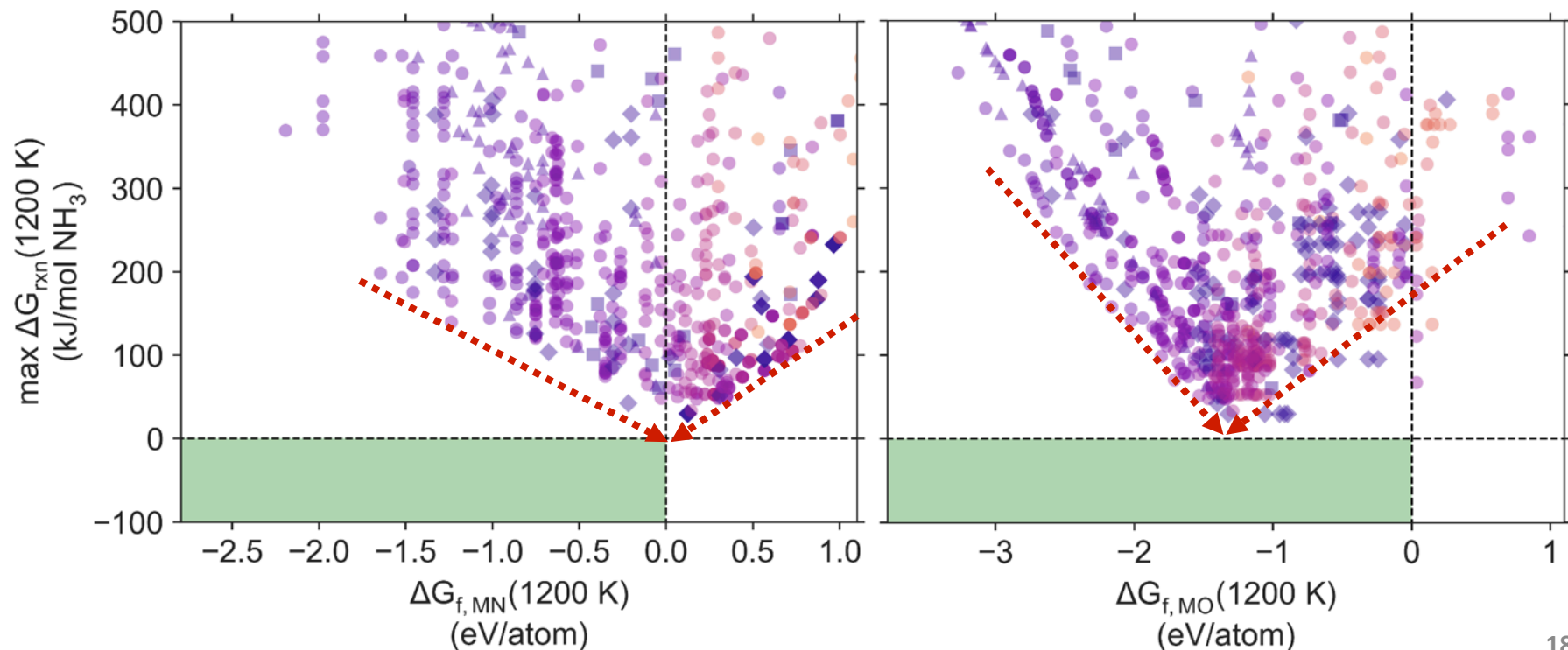
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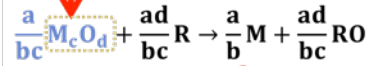
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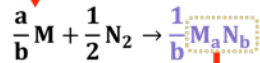
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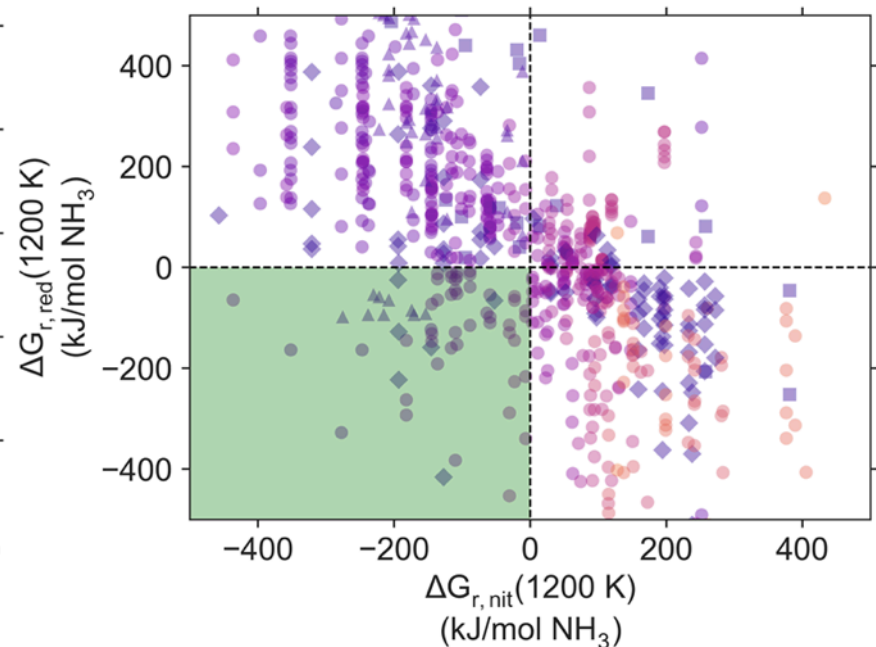
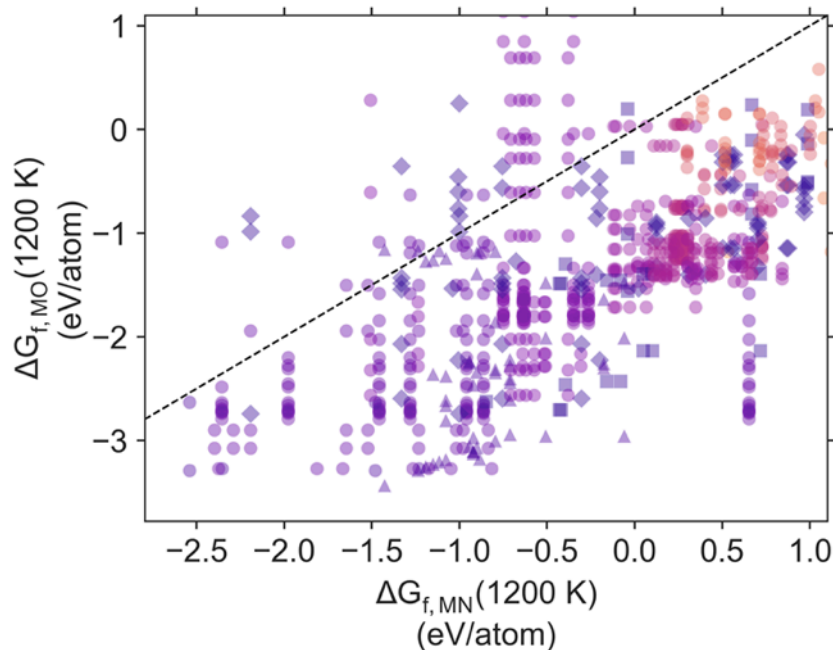
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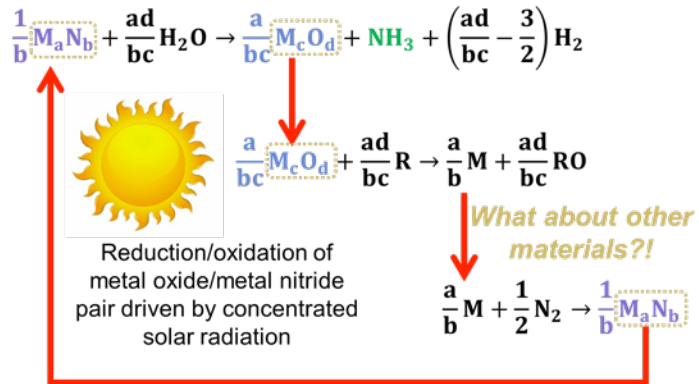
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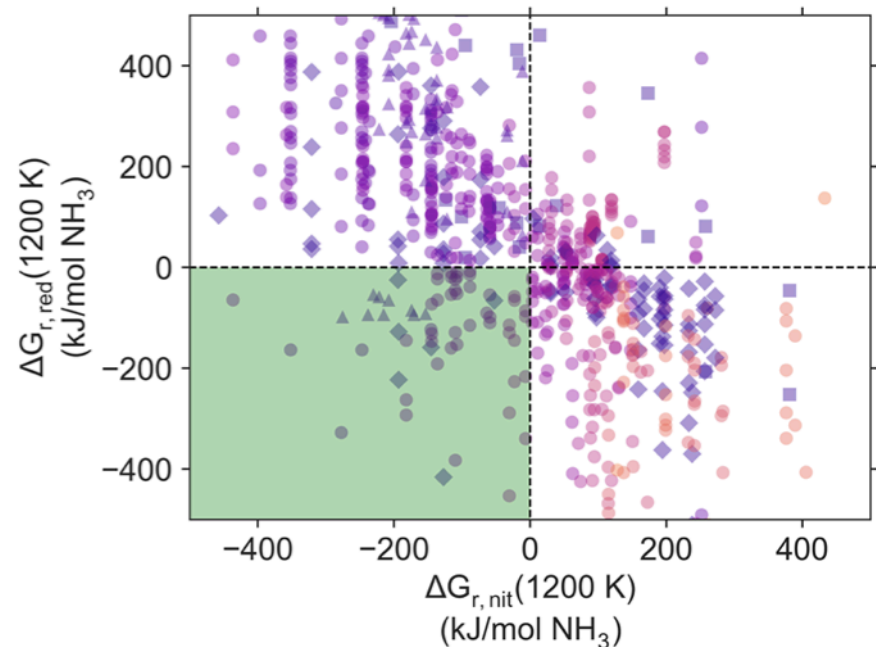
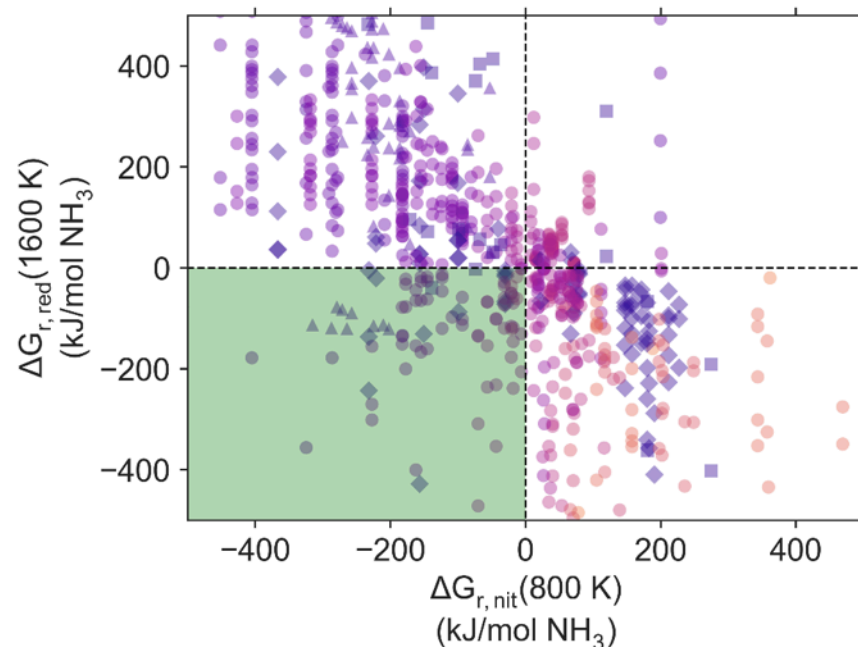
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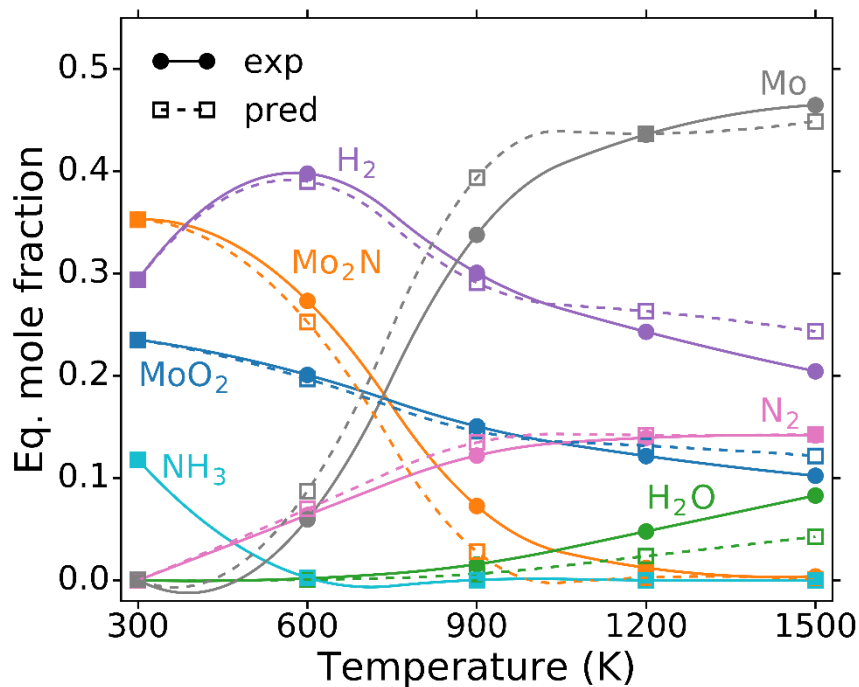
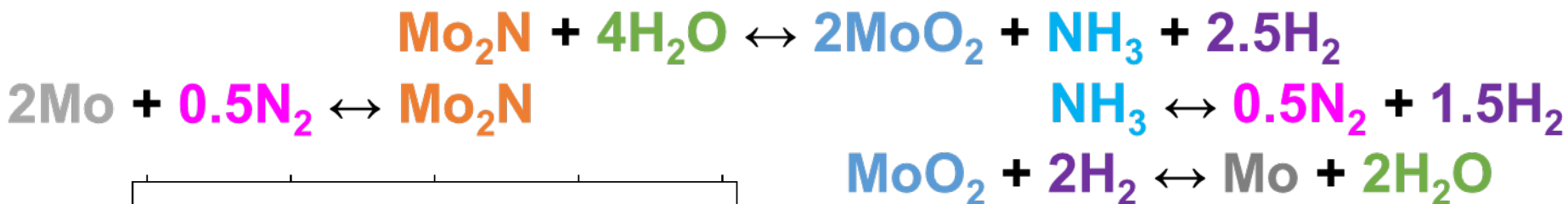
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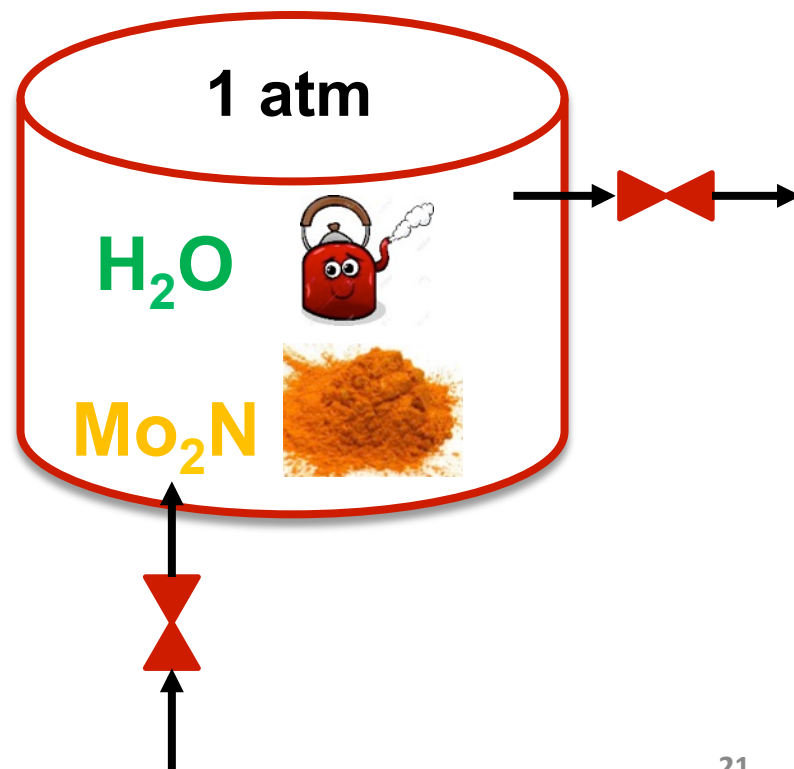
**Correlated stability of MN/
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What about going beyond reaction energetics?



Implicitly considers stability in addition to reaction energetics





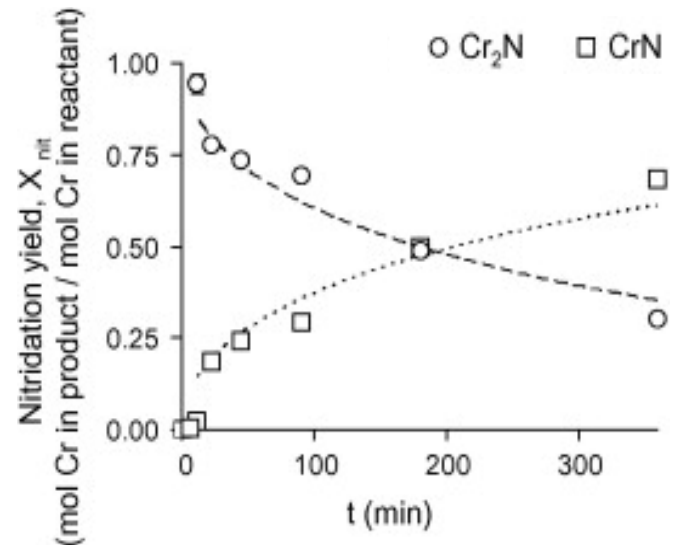
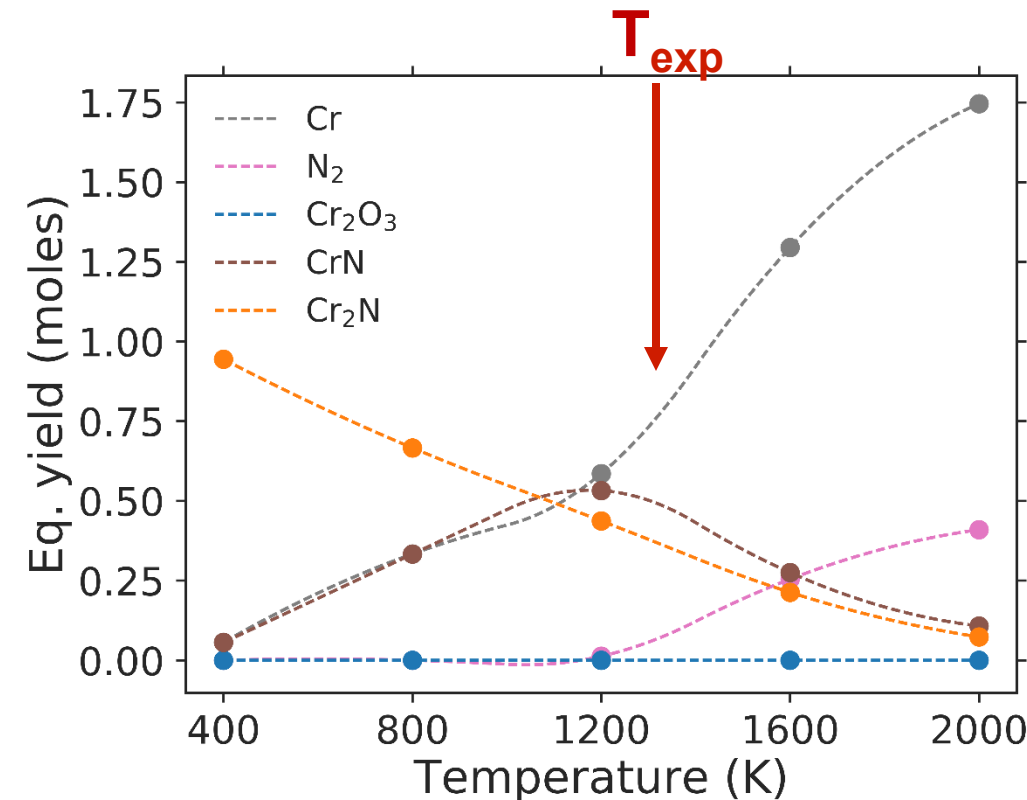
Cr

- $\text{Cr} + \text{N}_2 \rightarrow \text{Cr}_2\text{N}$ shown at 1000°C
- $\text{Cr}_2\text{N} + \text{H}_2\text{O} \rightarrow \text{Cr}_2\text{O}_3$ shown at 1000°C
- $\text{Cr}_2\text{O}_3 + \text{H}_2 \rightarrow \text{Cr}$ failed at 1600°C

Benchmarking equilibrium analysis

Cr

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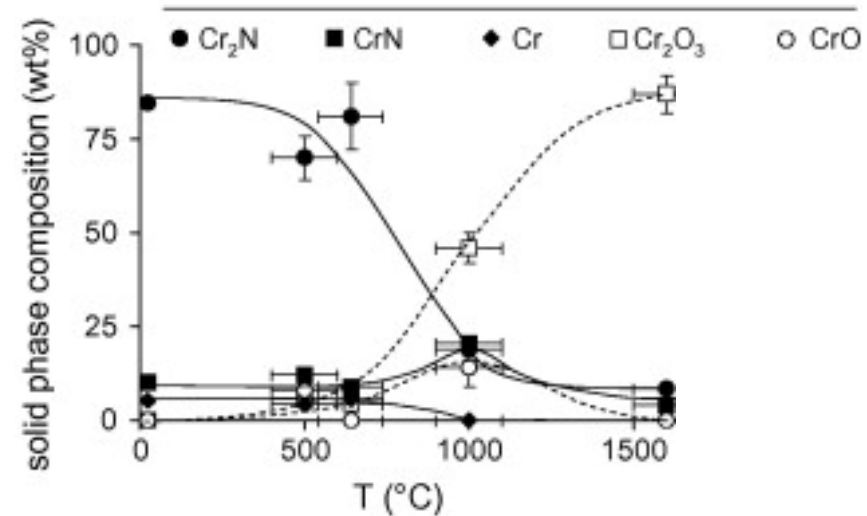
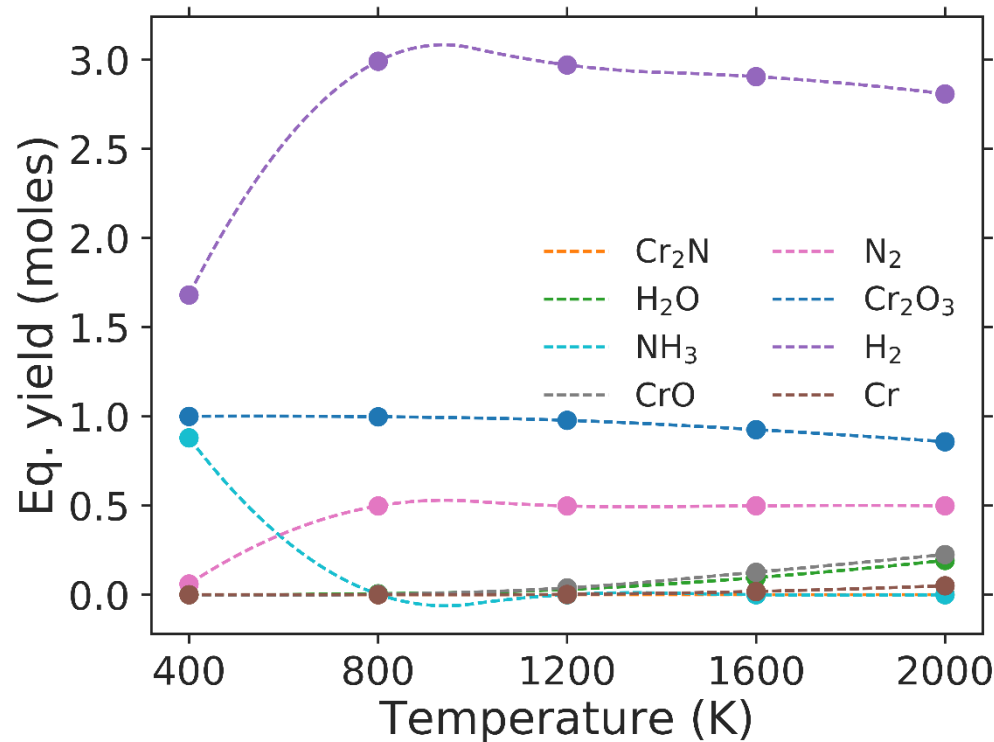


**Mix of $\text{CrN}/\text{Cr}_2\text{N}$
predicted and observed**

Benchmarking equilibrium analysis

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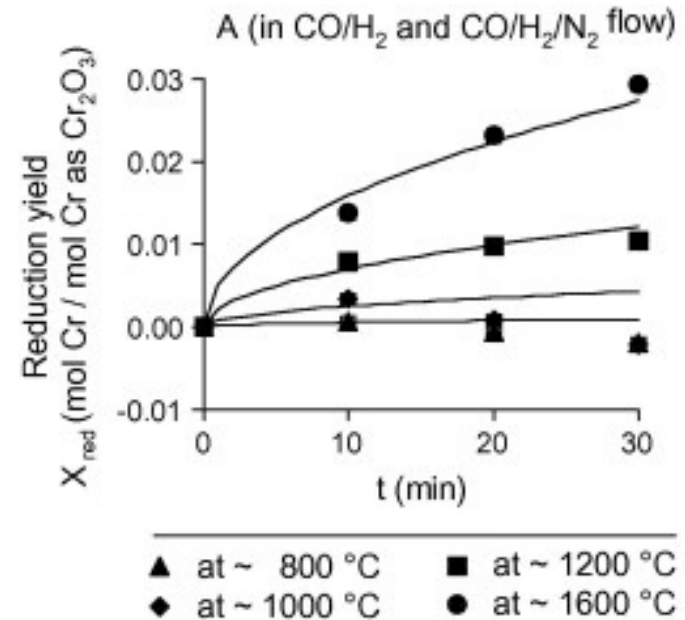
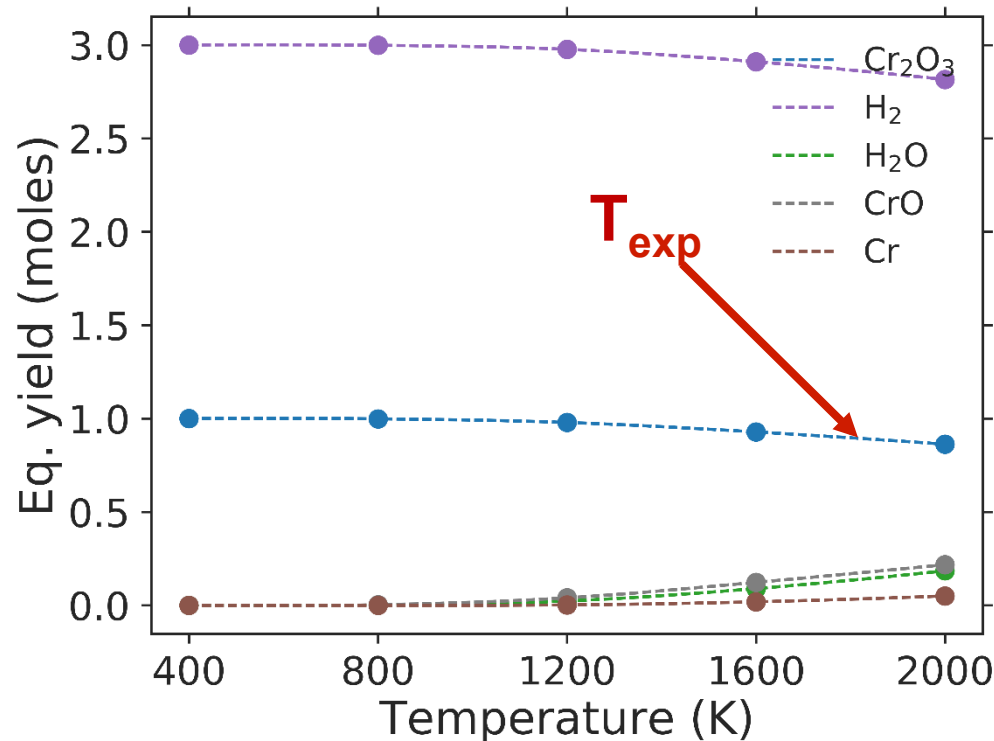


***Nitride corrosion
predicted and observed***

Benchmarking equilibrium analysis

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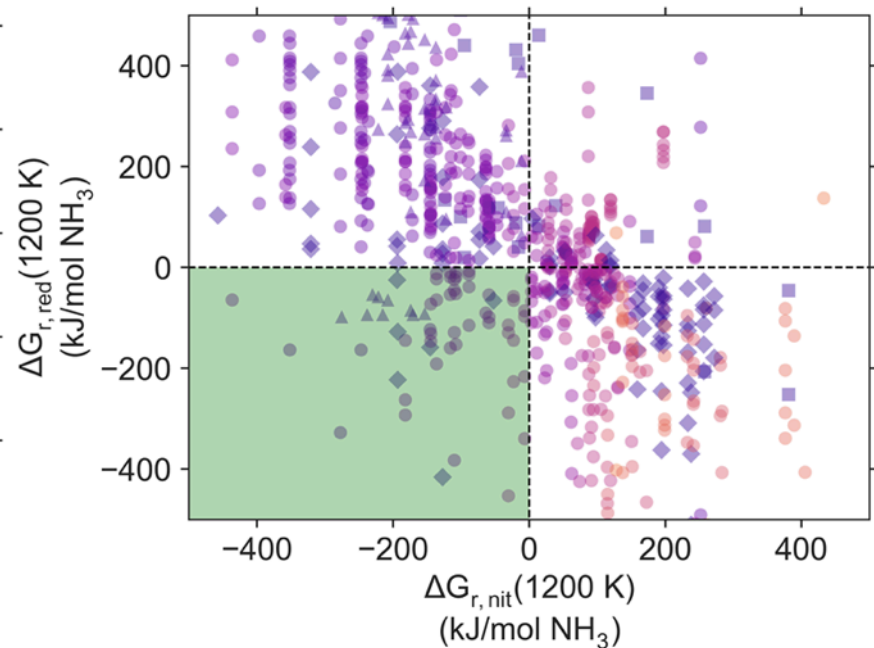
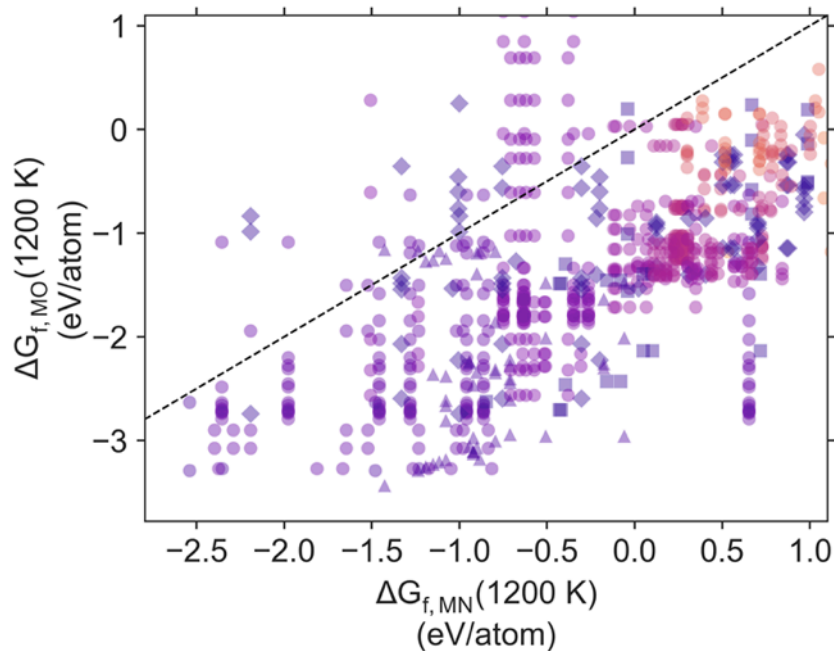


Minimal reduction yield predicted and observed

Systematic thermodynamic screening of MN/MO



✓ **STAS limited by correlated MO/MN stability**

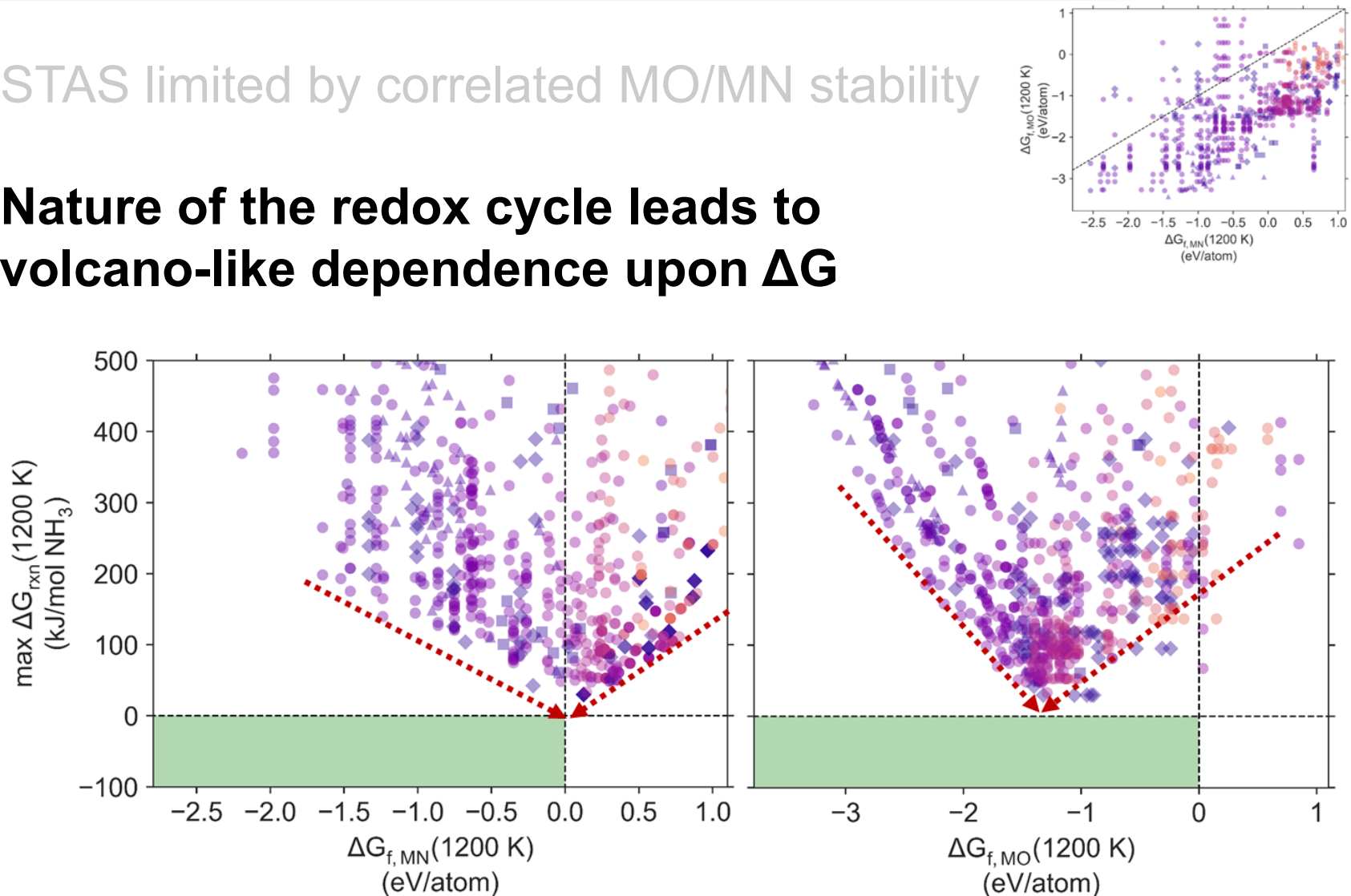


Systematic thermodynamic screening of MN/MO



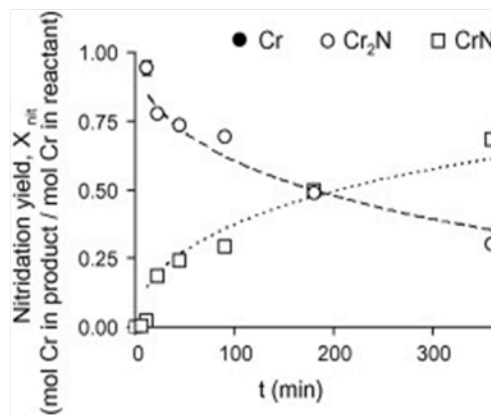
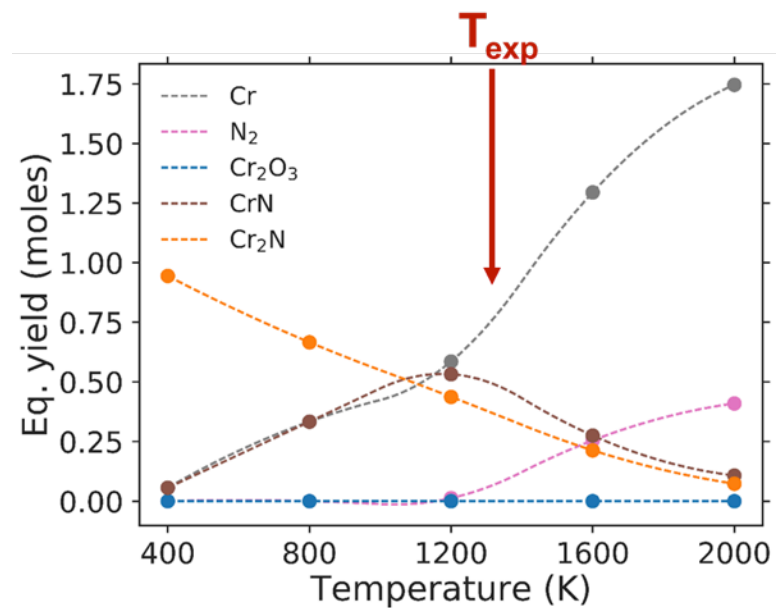
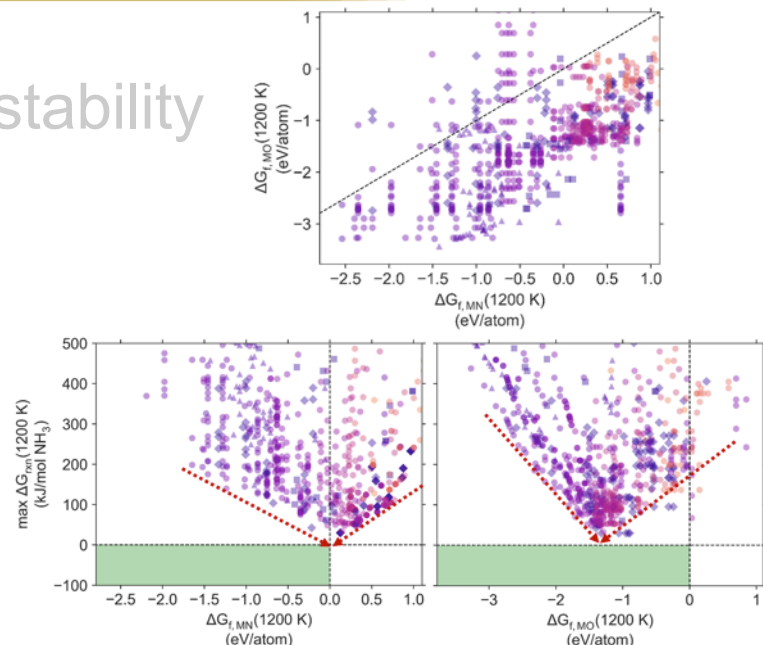
✓ STAS limited by correlated MO/MN stability

✓ **Nature of the redox cycle leads to volcano-like dependence upon ΔG**



Systematic thermodynamic screening of MN/MO

- ✓ STAS limited by correlated MO/MN stability
- ✓ Nature of the redox cycle leads to volcano-like dependence upon ΔG
- ✓ **Machine-learned eq analysis compares well with experiment**



**Mix of CrN/Cr₂N
predicted and observed**

Systematic thermodynamic screening of MN/MO



- ✓ STAS limited by correlated MO/MN stability
- ✓ Nature of the redox cycle leads to volcano-like dependence upon ΔG
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QUESTIONS?

